

Non-Relativistic Field Theory

What is the non-relativistic limit of the theory

$$\mathcal{L} = \partial_\mu \Phi^\dagger \partial^\mu \Phi - m^2 \Phi^\dagger \Phi - \lambda (\Phi^\dagger \Phi)^2 \quad ?$$

The equation of motion for $\lambda = 0$ is

$$(\partial^2 + m^2) \Phi = 0.$$

A mode with energy $E = m + \epsilon$ goes as $\Phi = e^{-iEt}$. So we let

$$\Phi(x, t) = e^{-imt} \phi(x, t)$$

$$\text{Then } \partial_0 \Phi(x, t) = -im e^{-imt} \phi(x, t) + e^{-imt} \partial_0 \phi(x, t)$$

$$\text{So } \partial_0^2 \Phi = \partial_0 e^{-imt} (-im \phi + \partial_0 \phi)$$

$$= [(-im)^2 - 2im \partial_0 \phi + \partial_0^2 \phi] e^{-imt}$$

and $(\partial^2 + m^2) \Phi = 0$ becomes

$$e^{-imt} [-m^2 - 2mi \partial_0 \phi + \partial_0^2 \phi] - \nabla^2 e^{-imt} \phi + m^2 e^{-imt} \phi = 0$$

↑ these cancel ↑

we drop $\partial_0^2 \phi$ compared to $-2mi \partial_0 \phi$, and

so we get

$$-2m i \partial_0 \phi - \nabla^2 \phi = 0$$

$$i \partial_0 \phi = - \frac{\nabla^2}{2m} \phi$$

which is Schrödinger's equation.

So to take the non-relativistic limit of a field theory, we replace

$$\tilde{\Phi}(x,t) = \frac{1}{\sqrt{2m}} e^{-imt} \phi(x,t)$$

so that

$$\frac{\partial \tilde{\Phi}^*}{\partial t} \frac{\partial \tilde{\Phi}}{\partial t} - m^2 \tilde{\Phi}^* \tilde{\Phi} \rightarrow \frac{1}{2m} \left\{ \left[(im + \frac{\partial}{\partial t}) \tilde{\Phi}^* \right] \left[(-im + \frac{\partial}{\partial t}) \tilde{\Phi} \right] - m^2 \tilde{\Phi}^* \tilde{\Phi} \right\}$$

$$\approx \frac{i}{2} \left(\tilde{\Phi}^* \frac{\partial \tilde{\Phi}}{\partial t} - \frac{\partial \tilde{\Phi}^*}{\partial t} \tilde{\Phi} \right)$$

Then \mathcal{L} becomes

$$\mathcal{L} = \partial \tilde{\Phi}^* \partial \tilde{\Phi} - m^2 \tilde{\Phi}^* \tilde{\Phi} - \lambda (\tilde{\Phi}^* \tilde{\Phi})^2$$

$$= \frac{i}{2} (\tilde{\Phi}^* \partial_0 \tilde{\Phi} - \partial_0 \tilde{\Phi}^* \tilde{\Phi}) - \frac{1}{2m} \partial_i \tilde{\Phi}^* \partial_i \tilde{\Phi} - \frac{\lambda}{4m^2} (\tilde{\Phi}^* \tilde{\Phi})^2$$

$$= i \tilde{\Phi}^* \partial_0 \tilde{\Phi} - \frac{1}{2m} \nabla \tilde{\Phi}^* \cdot \nabla \tilde{\Phi} - g^2 (\tilde{\Phi}^* \tilde{\Phi})^2 \quad \text{with } g^2 = \frac{\lambda}{4m^2}$$

The relativistic theory has a conserved current

$$J_m = i(\Phi^* \partial_m \Phi - \partial_m \Phi^* \Phi).$$

And $J_0 \approx \Phi^* \Phi$, while $J_i \approx \frac{\hbar}{2m} (\Phi^* \partial_i \Phi - \partial_i \Phi^* \Phi)$.

In the N.R theory,

$$\pi_\phi = \frac{\partial \mathcal{L}}{\partial \dot{\phi}} = i\phi^*$$

so $[\phi(x, t), \pi_\phi(x', t)] = i \delta(x - x')$

defines

$$[\phi(x, t), i\phi^*(x', t)] = i \delta(x - x')$$

or $[\phi(x, t), \phi^*(x', t)] = \delta(x - x')$.

In condensed-matter physics, there's often a mean density ρ , so it's useful to write

$$\phi = \sqrt{\rho} e^{i\theta} \quad \text{a / ker which}$$

$$\mathcal{L} = i\sqrt{\rho} e^{-i\theta} \partial_0 (\sqrt{\rho} e^{i\theta}) - \frac{1}{2m} [\nabla (\sqrt{\rho} e^{-i\theta}) \cdot \nabla (\sqrt{\rho} e^{i\theta})] - g^2 \rho^2$$

$$= \frac{i}{2} \dot{\rho} - \rho \dot{\theta} - \frac{1}{2m} \left[\rho (\nabla \theta)^2 + \frac{1}{4\rho} (\nabla \rho)^2 \right] - g^2 \rho^2$$

total divergence, which we can drop.

Now $\pi_\theta = \frac{\partial \mathcal{L}}{\partial \dot{\theta}} = -p$ so

$$[\theta(x, t), -p(x', t)] = i \delta(x - x')$$

so

$$[p(x, t), \theta(x', t)] = i \delta(x - x').$$

Integrate: $N = \int dx p(x, t)$ get

$$[N, \theta(x', t)] = i$$

Number is conjugate to phase.

Incidentally the physics of these two \mathcal{L} 's is the same:

$$\mathcal{L} = |\partial\phi|^2 - m^2 |\phi|^2 - \lambda |\phi|^4 \quad \text{and}$$

$$\mathcal{L} = |\partial\phi|^2 - m^2 |\phi|^2 + 2\sigma |\phi|^2 + \frac{1}{\lambda} \sigma^2$$

which some call a Hubbard-Stratonovich transformation.

The reason is that σ has no kinetic term and so its field equation is

$$0 = \frac{\partial \mathcal{L}}{\partial \sigma} = 2|\phi|^2 + \frac{2\sigma}{\lambda} = 0$$

$$\text{or } \sigma = -\lambda |\phi|^2.$$

Note that the field ϕ is repulsive.

$$L \sim -\lambda |\phi|^4 \quad \text{so} \quad H \sim \lambda |\phi|^4$$

Making piles of $|\phi|^2$ costs energy.

In condensed-matter physics, one often wants a finite density $\bar{\rho}$ in the system. So the N-R L is

$$L = i\phi^* \partial_t \phi - \frac{1}{2m} \nabla \phi^* \cdot \nabla \phi - g^2 (\phi^* \phi - \bar{\rho})^2.$$