

Spontaneous Emission in SI units

Let's compute the lifetime of an excited state of atomic hydrogen. If the transition is to go in the dipole approximation, then the selection rules

$l_f = l_i \pm 1$ $m_f = m_i$ or $m_f = m_i \pm 1$
 will apply. To have only one such transition, we will use the state $n=2, l=1, m=m_i$ as the initial state and $n=1, l=0, m=0$ plus one photon as the final state.

So we must compute

$$\langle 100k | S(t,0) | 21m \rangle = -\frac{i}{\hbar} \int_0^t dt' \langle 100k | e^{i(H_{0m} + H_{0f})t'/\hbar} \left(-\frac{q}{m} \vec{A}(\vec{x},0) \cdot \vec{p} \right) e^{-i(H_{0m} + H_{0f})t'/\hbar} | 21m \rangle \quad (1)$$

The field $A(\vec{x},0)$ is

$$\vec{A}(\vec{x},0) = \sum_{\vec{k}, \lambda} \left(\frac{\hbar}{2\epsilon_0 V \omega_k} \right)^{\frac{1}{2}} \left[\epsilon_{\vec{k}, \lambda}(\omega) a_{\vec{k}, \lambda}(\omega) e^{i\vec{k} \cdot \vec{x}} + \epsilon_{\vec{k}, \lambda}^*(\omega) a_{\vec{k}, \lambda}^\dagger(\omega) e^{-i\vec{k} \cdot \vec{x}} \right] \quad (2)$$

Now the photon part of $S(t,0)$ is just

$$\langle \vec{k}, \lambda | a_{\vec{k}, \lambda}^\dagger(\omega) | 0 \rangle = \delta_{\vec{k}, \vec{k}'} \delta_{\lambda, \lambda'} \quad (3)$$

So

$$\langle 100k | S(t,0) | 21m \rangle = \frac{iq}{\hbar m} \int_0^t e^{i(E_1 + \hbar\omega - E_2)t'/\hbar} \left(\frac{\hbar}{2\epsilon_0 V \omega_k} \right)^{\frac{1}{2}} \langle 100 | \epsilon_{\vec{k}, \lambda}^*(\omega) \cdot \vec{p} | 21m \rangle e^{-i\vec{k} \cdot \vec{x}} dt' \quad (4)$$

Doing the t' -integral, we get

$$\langle 100h | S(t,0) | 21m \rangle = \frac{q}{m} \left(\frac{t}{2\epsilon_0 V \omega_k} \right)^{\frac{1}{2}} \left(\frac{e^{i(E_1 + \hbar\omega - E_2)t/\hbar} - 1}{E_1 + \hbar\omega - E_2} \right) \langle 100 | \vec{\epsilon} \cdot \vec{p} e^{-i\vec{k} \cdot \vec{x}} | 21m \rangle. \quad (5)$$

So the probability $P(t)$ is

$$P(t) = |\langle 100h | S(t,0) | 21m \rangle|^2 \\ = \frac{q^2}{m^2} \frac{t}{2\epsilon_0 V \omega_k} \frac{4 \sin^2(E_1 + \hbar\omega - E_2)t/2\hbar}{(E_1 + \hbar\omega - E_2)^2} |\langle 100 | \vec{\epsilon} \cdot \vec{p} e^{-i\vec{k} \cdot \vec{x}} | 21m \rangle|^2. \quad (6)$$

Now our helpful δ -function formula is

$$\lim_{t \rightarrow \infty} \frac{\sin^2(E_1 + \hbar\omega - E_2)t/2\hbar}{(E_1 + \hbar\omega - E_2)^2} = \frac{\pi t}{2\hbar} \delta(E_1 + \hbar\omega - E_2). \quad (7)$$

So apart from a sum over final states, the transition rate is

$$\hat{W} = \frac{dP}{dt} = \frac{\pi q^2 t}{m^2 \epsilon_0 V \omega_k} |\langle 100 | \vec{\epsilon}_r(\vec{k}) \cdot \vec{p} e^{-i\vec{k} \cdot \vec{x}} | 21m \rangle|^2 \delta(E_1 + \hbar\omega - E_2). \quad (8)$$

Using $q^2 = \alpha \hbar c 4\pi \epsilon_0$ with $\alpha \approx 1/137.04$, we have

$$\hat{W} = \frac{4\pi^2 \alpha \hbar c}{m^2 V \omega_k} |\langle 100 | \vec{\epsilon}_r(\vec{k}) \cdot \vec{p} e^{-i\vec{k} \cdot \vec{x}} | 21m \rangle|^2 \delta(E_1 + \hbar\omega - E_2). \quad (9)$$

Recall now that $\vec{p} = \frac{m}{i\hbar} [\vec{x}, H_0] = \frac{m}{i\hbar} [\vec{x}, \frac{1}{2} m \vec{v}^2]$. (10)

So by making the dipole approximation

$$e^{-i\vec{k}\cdot\vec{x}} = 1 - i\vec{k}\cdot\vec{x} + \dots \approx 1 \quad (11)$$

which is valid to about 1 part in 1000, we find

$$\hat{W} = \frac{4\pi^2 \alpha \hbar c}{m^2 V \omega k} \left| \langle 100 | \frac{m}{i\hbar} [\vec{E}_n^{(*)}(\omega) \cdot \vec{x}, H_{0M}] | 21m \rangle \right|^2 \delta(E_1 + \hbar\omega - E_2) \quad (12)$$

$$= \frac{4\pi^2 \alpha c}{\hbar V \omega k} (E_2 - E_1)^2 \left| \langle 100 | \vec{E}_n^{(*)}(\omega) \cdot \vec{x} | 21m \rangle \right|^2 \delta(E_1 + \hbar\omega - E_2) \quad (13)$$

Since the δ -function keeps $E_2 - E_1 = \hbar\omega k$, we have

$$\hat{W} = \frac{4\pi^2 \alpha c}{V} (E_2 - E_1) \left| \langle 100 | \vec{E}_n^{(*)}(\omega) \cdot \vec{x} | 21m \rangle \right|^2 \delta(E_1 + \hbar\omega - E_2) \quad (14)$$

We may choose our coordinate system so that the magnetic quantum number $m=0$. In this case, only the z-component of \vec{x} survives

$$\langle 100 | x_i | 210 \rangle = \delta_{i3} \langle 100 | z | 210 \rangle, \quad (15)$$

Now we will now show that

$$\langle 100 | z | 210 \rangle = \frac{7}{2\sqrt{2}} a_0 = \frac{7}{3^5} \frac{\hbar}{\alpha m c} \quad (16)$$

in which $a_0 = \hbar / (\alpha m c) \cong 0.53 \text{ \AA}$ is the Bohr radius.

We use

$$R_{10}(r) = \frac{2}{a_0^{3/2}} e^{-r/a_0} \quad (a)$$

$$R_{21}(r) = \frac{1}{\sqrt{3}} \frac{1}{(2a_0)^{3/2}} \frac{r}{a_0} e^{-r/(2a_0)} \quad (b)$$

and $Y_0^0 = \frac{1}{\sqrt{4\pi}}$ and $Y_1^0(\theta, \phi) = \sqrt{\frac{3}{4\pi}} \cos\theta$. (c)

$$\begin{aligned} \langle 100 | Z | 210 \rangle &= \int d\Omega \int_0^\infty r^2 dr \frac{2}{a_0^{3/2}} \frac{e^{-r/a_0}}{\sqrt{4\pi}} r \cos\theta \frac{1}{\sqrt{3}} \frac{e^{-r/(2a_0)}}{(2a_0)^{3/2}} \frac{r}{a_0} \sqrt{\frac{3}{4\pi}} \cos\theta \end{aligned}$$

$$= \frac{2\pi}{4\pi} a_0^4 \sqrt{2} \int_{-1}^1 d\cos\theta \int_0^\infty dr r^4 e^{-3r/2a_0} \cos^2\theta$$

$$= \frac{1}{3\sqrt{2}} \frac{1}{a_0^4} \int_0^\infty dr r^4 e^{-3r/2a_0}$$

$$= \frac{1}{3\sqrt{2}} \frac{1}{a_0^4} \frac{4!}{\left(\frac{3}{2a_0}\right)^5} = \frac{4 \cdot 3 \cdot 2 \cdot 2^5}{3^6 \sqrt{2}} = \frac{2^7}{3^5} \sqrt{2} a_0$$

$$= \frac{2^7}{3^5} \sqrt{2} \frac{\hbar}{\alpha m c} \quad (d)$$

Using this value for the matrix element
 $\langle 100 | z | 210 \rangle = 2^7 \sqrt{2} \hbar / (3^5 \alpha m c)$

we find for the transition rate

$$\hat{W} = \frac{4\pi^2 \alpha c}{V} (E_2 - E_1) |\epsilon_n(\mathbf{k})_3|^2 \frac{2^{15} \hbar^2}{3^{10} \alpha^2 m^2 c^2} \delta(E_1 + \hbar\omega - E_2). \quad (17)$$

in which $\epsilon_n(\mathbf{k})_3 = \epsilon_n(\mathbf{k}) \cdot \hat{\mathbf{z}}$ is the 3-component of the polarization vector $\vec{\epsilon}_n(\mathbf{k})$.

So

$$\hat{W} = \frac{2^{17} \pi^2 \hbar^2}{3^{10} \alpha c m^2} \frac{(E_2 - E_1)}{V} |\epsilon_n(\mathbf{k})_3|^2 \delta(E_1 + \hbar\omega - E_2). \quad (18)$$

The polarization vectors $\epsilon_1(\mathbf{k})$ and $\epsilon_2(\mathbf{k})$ span the subspace of \mathbb{R}^3 orthogonal to the vector \vec{k} . Thus

$$\sum_{n=1}^2 \epsilon_n(\mathbf{k}) \epsilon_n^\dagger(\mathbf{k}) = 1 - \hat{\mathbf{k}} \hat{\mathbf{k}}^T. \quad (19)$$

To sum \hat{W} over n , we will use

$$\begin{aligned} \sum_{n=1}^2 |\epsilon_n(\mathbf{k})_3|^2 &= \left(\sum_{n=1}^2 \epsilon_n(\mathbf{k}) \epsilon_n^\dagger(\mathbf{k}) \right)_{33} \\ &= \left(1 - \hat{\mathbf{k}} \hat{\mathbf{k}}^T \right)_{33} = \delta_{33} - \hat{k}_3 \hat{k}_3 \\ &= 1 - \hat{k}_3^2. \end{aligned} \quad (20)$$

If the vector \vec{k} has polar angles θ and ϕ , then

$$k_3 = \cos \theta \quad (21)$$

and the sum over polarizations is

$$\sum_{r=1}^2 |\epsilon_r(\vec{k})_3|^2 = 1 - \cos^2 \theta. \quad (22)$$

We now must sum over the momenta and spins of the emitted photon. Since

$$\vec{k} = \frac{2\pi}{L} \vec{n} \quad (23)$$

and

$$\vec{n} = \frac{L}{2\pi} \vec{k} \quad (24)$$

we may approximate the sum as the integral

$$\sum_{\vec{n}} = \left(\frac{L}{2\pi}\right)^3 \int d^3k = V \int \frac{d^3k}{(2\pi)^3} \quad (25)$$

$$\begin{aligned} \omega &= \frac{2^{17}}{3^{10}} \frac{\pi^2 \hbar^2 (E_2 - E_1)}{\alpha c m^2} \int \frac{d^3k}{(2\pi)^3} (1 - \cos^2 \theta) \delta(E_1 + \hbar\omega - E_2) \\ &= \frac{2^{17}}{3^{10}} \frac{\pi^2 \hbar^2 (E_2 - E_1)}{\alpha c m^2} \int_0^\infty \frac{k^2 d\hbar\omega}{\hbar c} \int_{-1}^1 dx \frac{(1-x^2)}{(2\pi)^2} \delta(E_1 + \hbar\omega - E_2) \\ &= \frac{2^{15}}{3^{10}} \frac{\hbar (E_2 - E_1)}{\alpha c^2 m^2} \frac{(E_2 - E_1)^2}{\hbar^2 c^2} \int_{-1}^1 dx (1-x^2). \quad (26) \end{aligned}$$

Since $\int_0^1 dx (1-x^2) = 2 - \left[\frac{x^3}{3} \right]_0^1 = 2 \left(1 - \frac{1}{3}\right) = \frac{4}{3}$ (27)

the rate is

$$W = \frac{2^{17}}{3^{11}} \frac{(E_2 - E_1)^3}{\alpha^3 h c^4 m^2} \quad (28)$$

Now in atomic hydrogen

$$E_n = -\frac{1}{2} m c^2 \frac{\alpha^2}{n^2} \quad (29)$$

So

$$E_2 - E_1 = \frac{1}{2} m c^2 \alpha^2 \left(1 - \frac{1}{4}\right) = \frac{3}{8} m c^2 \alpha^2 \quad (30)$$

So the rate is

$$W = \frac{2^{17}}{3^{11}} \frac{3^3}{2^9} \frac{m^3 c^6 \alpha^6}{\alpha^3 h c^4 m^2} \quad (31)$$

$$= \left(\frac{2}{3}\right)^8 \frac{m c^2 \alpha^5}{h} \quad (31)$$

The lifetime $\tau = 1/W$ then is

$$\tau = \frac{1}{W} = \left(\frac{3}{2}\right)^8 \frac{h}{m c^2} \alpha^{-5} \quad (32)$$

Numerically $mc^2 = 0.511 \text{ MeV}$

$$\hbar = 6.582 \times 10^{-22} \text{ MeV s}$$

$$\alpha = 1/137.036 \quad (33)$$

so the lifetime of the 2p state of atomic hydrogen is

$$\begin{aligned} \tau &= \left(\frac{3}{2}\right)^8 \frac{6.582 \times 10^{-22}}{0.511} (137.036)^5 \text{ s} \\ &= 1.6 \times 10^{-9} \text{ s} = 1.60 \text{ ns}. \end{aligned} \quad (34)$$

This 2p → 1s transition is said to be an "allowed transition" because it proceeds through the first term in

$$e^{ik \cdot x} = 1 + ik \cdot x + \dots, \quad (35)$$

They have lifetimes of nanoseconds. The 2s → 1s transition must use the $k \cdot x$ term which is smaller by a factor of about 10^{-3} . When squared, this suppression is 10^{-6} . Such transitions take milliseconds. They are called "magnetic-dipole" or "electric-quadrupole" transitions according to whether the anti-symmetric or symmetric part of $e^{ik \cdot x}$ causes the decay of the excited state.