

our formula (7.469) for  $f'(x)$  gives

$$f'(a) = \frac{u'(a)}{A} \int_a^b v(y) g(y) dy = 0 = f'(b) = \frac{v'(b)}{A} \int_a^b u(y) g(y) dy. \quad (7.473)$$

For instance, to solve the equation  $-f''(x) - f(x) = \exp x$ , with the mixed boundary conditions  $f(-\pi) = 0$  and  $f'(\pi) = 0$ , we choose from among the solutions  $\alpha \cos x + \beta \sin x$  of the homogeneous equation  $-f'' - f = 0$ , the functions  $u(x) = \sin x$  and  $v(x) = \cos x$ . Substituting them into the formula (7.468) and setting  $p(x) = 1$  and  $A = -W(x_0) = \sin^2(x_0) + \cos^2(x_0) = 1$ , we find as the Green's function

$$G(x, y) = \theta(x - y) \sin y \cos x + \theta(y - x) \sin x \cos y. \quad (7.474)$$

The solution  $f(x) = \int_{-\pi}^{\pi} G(x, y) e^y dy$  then is

$$\begin{aligned} f(x) &= \int_{-\pi}^{\pi} [\theta(x - y) \sin y \cos x + \theta(y - x) \sin x \cos y] e^y dy \\ &= -\frac{1}{2} (e^{-\pi} \cos x + e^{\pi} \sin x + e^x). \end{aligned} \quad (7.475)$$

□

#### 7.44 Principle of Stationary Action in Field Theory

If  $\phi(x)$  is a scalar field, and  $L(\phi)$  is its action density, then its action  $S[\phi]$  is the integral over all of spacetime

$$S[\phi] = \int L(\phi(x), \partial_a \phi(x)) d^4x. \quad (7.476)$$

The principle of least (or stationary) action says that the field  $\phi(x)$  that satisfies the classical equation of motion is the one for which the first-order change in the action due to any tiny variation  $\delta\phi(x)$  in the field vanishes,  $\delta S[\phi] = 0$ . To keep things simple, we'll assume that the action (or Lagrange) density  $L(\phi)$  is a function only of the field  $\phi$  and its first derivatives  $\partial_a \phi = \partial\phi/\partial x^a$ . The first-order change in the action then is

$$\delta S[\phi] = \int \left[ \frac{\partial L}{\partial \phi} \delta\phi + \frac{\partial L}{\partial (\partial_a \phi)} \delta(\partial_a \phi) \right] d^4x = 0 \quad (7.477)$$

in which we sum over the repeated index  $a$  from 0 to 3. Now  $\delta(\partial_a \phi) = \partial_a(\phi + \delta\phi) - \partial_a \phi = \partial_a \delta\phi$ . So we may integrate by parts and drop the surface

terms because we set  $\delta\phi = 0$  on the surface at infinity

$$\delta S[\phi] = \int \left[ \frac{\partial L}{\partial \phi} \delta\phi + \frac{\partial L}{\partial(\partial_a \phi)} \partial_a(\delta\phi) \right] d^4x = \int \left[ \frac{\partial L}{\partial \phi} - \partial_a \frac{\partial L}{\partial(\partial_a \phi)} \right] \delta\phi d^4x.$$

This first-order variation is zero for arbitrary  $\delta\phi$  only if the field  $\phi(x)$  satisfies Lagrange's equation

$$\partial_a \left( \frac{\partial L}{\partial(\partial_a \phi)} \right) \equiv \frac{\partial}{\partial x^a} \left[ \frac{\partial L}{\partial(\partial\phi/\partial x^a)} \right] = \frac{\partial L}{\partial \phi} \quad (7.478)$$

which is the classical equation of motion.

**Example 7.66** (Theory of a scalar field) The action density of a single scalar field  $\phi$  of mass  $m$  is  $L = \frac{1}{2} (\dot{\phi})^2 - \frac{1}{2} (\nabla\phi)^2 - \frac{1}{2} m^2 \phi^2$  or equivalently  $L = -\frac{1}{2} \partial_a \phi \partial^a \phi - \frac{1}{2} m^2 \phi^2$ . Lagrange's equation (7.478) is then

$$\nabla^2 \phi - \ddot{\phi} = \partial_a \partial^a \phi = m^2 \phi \quad (7.479)$$

which is the Klein-Gordon equation (7.41).  $\square$

In a theory of several fields  $\phi_1, \dots, \phi_n$  with action density  $L(\phi_k, \partial_a \phi_k)$ , the fields obey  $n$  copies of Lagrange's equation one for each field  $\phi_k$

$$\partial_a \left( \frac{\partial L}{\partial(\partial_a \phi_k)} \right) = \frac{\partial}{\partial x^a} \left( \frac{\partial L}{\partial(\partial_a \phi_k)} \right) = \frac{\partial L}{\partial \phi_k}. \quad (7.480)$$

Application of the principle of stationary action to the action  $\int R\sqrt{g} d^4x$  gives Einstein's equations as shown in section 13.37.

### 7.45 Symmetries and Conserved Quantities in Field Theory

A transformation of the coordinates  $x^a$  or of the fields  $\phi_i$  and their derivatives  $\partial_a \phi_i$  that leaves the action density  $L(\phi_i, \partial_a \phi_i)$  invariant is a **symmetry** of the theory. Such a symmetry implies that something is conserved or time independent.

Suppose that due to a symmetry a Lagrange density  $L(\phi_i, \partial_a \phi_i)$  is unchanged when the fields  $\phi_i$  and their derivatives  $\partial_a \phi_i$  change by  $\delta\phi_i$  and by  $\delta(\partial_a \phi_i) = \partial_a(\delta\phi_i)$  which need not be small:

$$0 = \delta L = \sum_i \frac{\partial L}{\partial \phi_i} \delta\phi_i + \frac{\partial L}{\partial \partial_a \phi_i} \partial_a \delta\phi_i. \quad (7.481)$$

Then using Lagrange's equations (7.480), we can rewrite  $\partial L/\partial\phi_i$  and get

$$0 = \sum_i \left( \partial_a \frac{\partial L}{\partial \partial_a \phi_i} \right) \delta\phi_i + \frac{\partial L}{\partial \partial_a \phi_i} \partial_a \delta\phi_i = \partial_a \sum_i \frac{\partial L}{\partial \partial_a \phi_i} \delta\phi_i \quad (7.482)$$

which says that the **current**

$$J^a = \sum_i \frac{\partial L}{\partial \partial_a \phi_i} \delta \phi_i \quad (7.483)$$

has zero divergence,  $\partial_a J^a = 0$ . Thus the time derivative of the volume integral of the charge density  $J^0$

$$Q_V = \int_V J^0 d^3x \quad (7.484)$$

is the flux of current  $\mathbf{J}$  entering through the boundary  $S$  of the volume  $V$

$$\dot{Q}_V = \int_V \partial_0 J^0 d^3x = - \int_V \nabla \cdot \mathbf{J} d^3x = - \int_S \mathbf{J} \cdot d^2\mathbf{S}. \quad (7.485)$$

If no current enters  $V$ , then the charge  $Q$  inside  $V$  is conserved. When the volume  $V$  is the whole universe, the charge is the integral over all of space

$$Q = \int J^0 d^3x = \int \sum_i \frac{\partial L}{\partial \dot{\phi}_i} \delta \phi_i d^3x = \int \sum_i \pi_i \delta \phi_i d^3x \quad (7.486)$$

in which  $\pi_i$  is the momentum conjugate to the field  $\phi_i$

$$\pi_i = \frac{\partial L}{\partial \dot{\phi}_i}. \quad (7.487)$$

**Example 7.67** ( $O(n)$  symmetry and its charge) Suppose the action density  $L$  is the sum of  $n$  copies of the quadratic action density

$$L = \sum_{i=1}^n \frac{1}{2} (\dot{\phi}_i)^2 - \frac{1}{2} (\nabla \phi_i)^2 - \frac{1}{2} m^2 \phi_i^2 = -\frac{1}{2} \partial_a \phi \partial^a \phi - \frac{1}{2} m^2 \phi^2. \quad (7.488)$$

Let  $A_{ij}$  be any constant antisymmetric matrix,  $A_{ij} = -A_{ji}$ . Then if the fields change by  $\delta \phi_i = \epsilon \sum_j A_{ij} \phi_j$ , the change (7.481) in the action density

$$\delta L = -\epsilon \sum_{i,j=1}^n \partial^a \phi_i A_{ij} \partial_a \phi_j + m^2 \phi_i A_{ij} \phi_j = 0 \quad (7.489)$$

vanishes. Thus the charge (7.486) associated with the matrix  $A$

$$Q_A = \int \sum_{i,j} \pi_i \delta \phi_i d^3x = \epsilon \int \sum_{i,j} \pi_i A_{ij} \phi_j d^3x \quad (7.490)$$

is conserved. There are  $n(n-1)/2$  antisymmetric  $n \times n$  imaginary matrices; they generate the group  $O(n)$  of  $n \times n$  orthogonal matrices (example 11.3).

□

If an action density  $L(\phi_i, \partial_a \phi_i)$  depends upon the spacetime coordinate  $x^a$  only through the fields  $\phi_i$  and their derivatives  $\partial_a \phi_i$ , then its spacetime derivative is

$$\frac{\partial L}{\partial x^a} = \sum_i \left( \frac{\partial L}{\partial \phi_i} \frac{\partial \phi_i}{\partial x^a} + \frac{\partial L}{\partial \partial_b \phi_i} \frac{\partial^2 \phi_i}{\partial x^b \partial x^a} \right). \quad (7.491)$$

Using Lagrange's equations (7.480) to rewrite  $\partial L / \partial \phi_i$ , we find

$$\begin{aligned} 0 &= \left( \sum_i \partial_b \left( \frac{\partial L}{\partial \partial_b \phi_i} \right) \partial_a \phi_i + \frac{\partial L}{\partial \partial_b \phi_i} \frac{\partial^2 \phi_i}{\partial x^b \partial x^a} \right) - \frac{\partial L}{\partial x^a} \\ 0 &= \partial_b \left[ \left( \sum_i \frac{\partial L}{\partial \partial_b \phi_i} \frac{\partial \phi_i}{\partial x^a} \right) - \delta_a^b L \right] \end{aligned} \quad (7.492)$$

which says that the **energy-momentum tensor** of the theory

$$T_a^b = \sum_i \frac{\partial L}{\partial \partial_b \phi_i} \frac{\partial \phi_i}{\partial x^a} - \delta_a^b L \quad (7.493)$$

has zero divergence,  $\partial_b T_a^b = 0$ .

Thus the time derivative of the 4-momentum  $P_{aV}$  inside a volume  $V$

$$P_{aV} = \int_V \left( \sum_i \frac{\partial L}{\partial \partial_0 \phi_i} \frac{\partial \phi_i}{\partial x^a} - \delta_a^0 L \right) d^3x = \int_V T_a^0 d^3x \quad (7.494)$$

is equal to the flux entering through  $V$ 's boundary  $S$

$$\partial_0 P_{aV} = \int_V \partial_0 T_a^0 d^3x = - \int_V \partial_k T_a^k d^3x = - \int_S T_a^k d^2S_k. \quad (7.495)$$

And so if an action density has no explicit dependence on any of the spacetime coordinates  $x$ , then its **energy**  $P_0$  and **momentum**  $\vec{P}$  are conserved.

**Example 7.68** (Conservation of angular momentum) Under an infinitesimal rotation by  $\boldsymbol{\theta}$ , spatial coordinates change by  $\delta \mathbf{x} = \boldsymbol{\theta} \times \mathbf{x}$  and the change in a scalar field is  $\delta \phi(t, \mathbf{x}) = \nabla \phi \cdot \delta \mathbf{x} = \nabla \phi \cdot (\boldsymbol{\theta} \times \mathbf{x}) = \boldsymbol{\theta} \cdot (\mathbf{x} \times \nabla \phi)$ . So if the spatial integral of the Lagrange density  $L$  is unchanged under rotations

$$0 = \int d^3x \frac{\partial L}{\partial \phi} \delta \phi + \frac{\partial L}{\partial \partial_a \phi} \partial_a \delta \phi = \int d^3x \partial_a \frac{\partial L}{\partial \partial_a \phi} \delta \phi \quad (7.496)$$

then the time derivative of the angular momentum

$$\mathbf{J} = \int d^3x \frac{\partial L}{\partial \dot{\phi}} \delta \mathbf{x} = \int d^3x \frac{\partial L}{\partial \dot{\phi}} (\mathbf{x} \times \nabla \phi) \quad (7.497)$$

vanishes as long as the fields fall to zero as  $|\mathbf{x}| \rightarrow \infty$ . If a rotation changes

the field component  $\phi^\ell$  by  $\delta\phi^\ell = \boldsymbol{\theta} \cdot (\mathbf{x} \times \nabla\phi^\ell - i\mathbf{S}_m^\ell\phi^m)$ , where  $S$  represents the spin of the field, then the conserved angular momentum is

$$\mathbf{J} = \int d^3x \frac{\partial L}{\partial \dot{\phi}^\ell} (\mathbf{x} \times \nabla\phi^\ell - i\mathbf{S}_m^\ell\phi^m). \tag{7.498}$$

□

The momentum  $\pi_i(x)$  that is canonically conjugate to the field  $\phi_i(x)$  is the derivative of the action density  $L$  with respect to the time derivative of the field

$$\pi_i = \frac{\partial L}{\partial \dot{\phi}_i}. \tag{7.499}$$

If one can express the time derivatives  $\dot{\phi}_i$  of the fields in terms of the fields  $\phi_i$  and their momenta  $\pi_i$ , then the hamiltonian of the theory is the spatial integral of

$$H = P_0 = T_0^0 = \left( \sum_{i=1}^n \pi_i \dot{\phi}_i \right) - L \tag{7.500}$$

in which  $\dot{\phi}_i = \dot{\phi}_i(\phi, \pi)$ .

**Example 7.69** (Hamiltonian of a scalar field) For the lagrangian  $L$  of example 7.66, the hamiltonian density (7.500) is  $H = \frac{1}{2} \pi^2 + \frac{1}{2} (\nabla\phi)^2 + \frac{1}{2} m^2\phi^2$ . □

**Example 7.70** (Euler’s theorem and the Nambu-Goto string) When the action density is a first-degree homogeneous function (section 7.10) of the time derivatives of the fields, as is that of the Nambu-Goto string

$$L = -\frac{T_0}{c} \sqrt{(\dot{X} \cdot X')^2 - (\dot{X})^2 (X')^2}, \tag{7.501}$$

Euler’s theorem (7.112) implies that the energy density (7.500) vanishes identically, independently of the equations of motion,

$$E^0 = \frac{\partial L}{\partial \dot{X}^\mu} \dot{X}^\mu - L = 0. \tag{7.502}$$

□