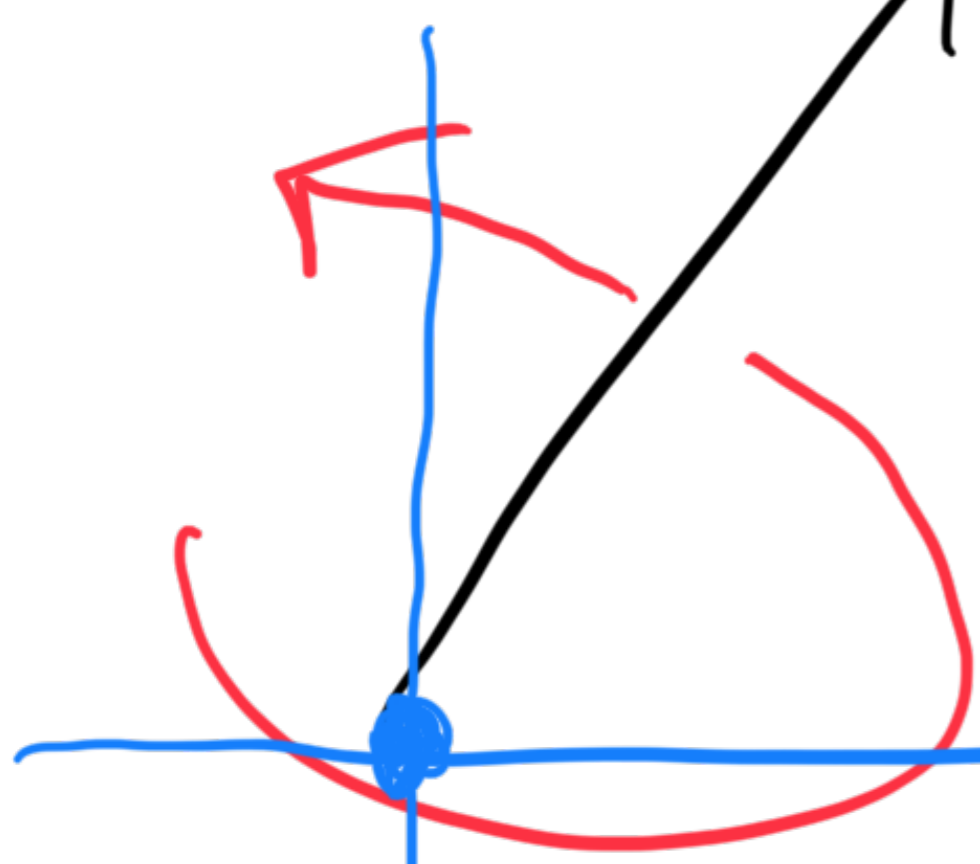


# 500 Seminar

N  
R  
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axis of notation

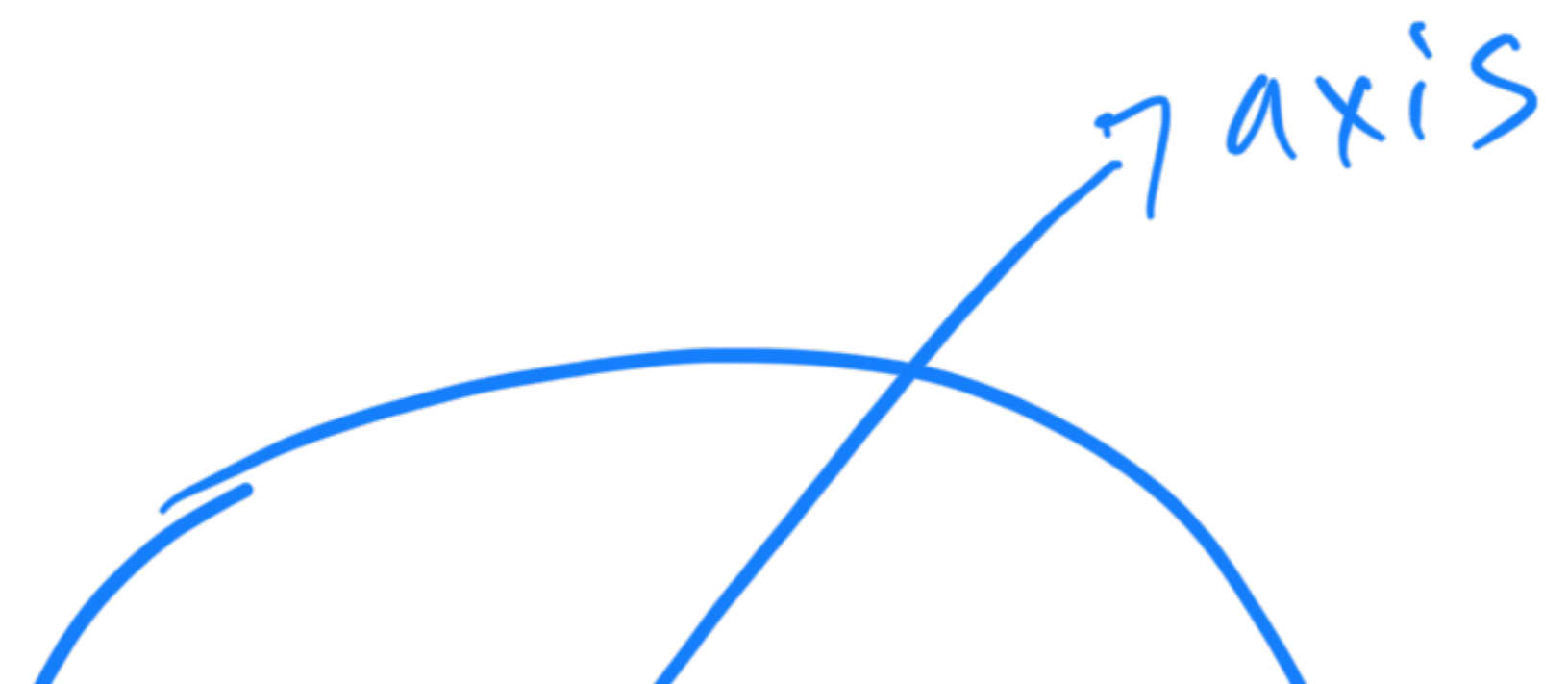


$$d\vec{r} = d\theta \times \vec{r}$$

$$\Delta \vec{r} = \vec{\theta} \times \vec{r}$$

$\hat{\theta}$  axis

$|\vec{\theta}|$  radians in notation





$\hat{n}_i$  inertial, fixed unit vectors  
 $\hat{e}_i(t)$  unit vectors rotating with Earth

$$\vec{r}(t) = r_i \hat{n}_i = r_i(t) \hat{n}_i$$

$$\vec{v}(t) = v_{e_i}(t) \hat{e}_i(t)$$

$$\dot{\vec{r}}(t) = \vec{v}(t) = \dot{r}_i(t) \hat{n}_i$$

$$\vec{r}(t) = r_e(t) = r_{e_i}(t) \hat{e}_i(t)$$

$$\dot{\vec{r}}(t) = \dot{r}_{e_i}(t) \hat{e}_i(t) + r_{e_i}(t) \dot{\hat{e}}_i(t)$$

$$\dot{\hat{e}}_i(t) = \dot{\theta} \times \hat{e}_i$$

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$$\Delta \vec{r} = \vec{\theta} \times \vec{r} \quad \dot{\theta} = \omega$$

$$\dot{\vec{r}} = \omega \times \vec{r} + \theta \times \vec{v}$$

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$\vec{r}'_i(t) = \hat{e}_i(t) r'_i$   $\vec{r}'_i$  notation system

$\dot{\vec{r}} = \dot{\vec{r}}' = \dot{\hat{e}}_i r'_i + \hat{e}_i \dot{r}'_i$

$$\dot{\hat{e}}_i = \omega \times \hat{e}_i \quad \omega = \dot{\theta}$$

$$\dot{\vec{r}} = \dot{\vec{r}}' = (\omega \times \hat{e}_i) r'_i + \hat{e}_i \dot{r}'_i$$

$$v = \omega \times r + v'$$

$$v_f = \omega \times r + v'$$

$$= \omega \times r + v'$$

$$\frac{d}{dt} = \omega \times + \frac{d}{dt_r}$$

real

apparent  
time derivative

time  
derivative

$$a = \frac{d^2 r}{dt^2} = \left( \omega \times + \frac{d}{dt_r} \right) \left( \omega \times + \frac{d}{dt_r} \right) r$$

real

$$r = r + \omega \times r + \dots$$

-2,

$$= \omega \times (\omega \times r) + \dots$$

$$\frac{d}{dt} (\omega \times r) + \frac{d^2 r}{dt^2}$$

$a$   
Coriolis

$$\vec{a}_f = \vec{a}' + \omega \times (\omega \times \vec{r}) + 2\omega \times \vec{v}'$$

centrifugal acceleration

Euler's

$$\vec{a}' = \vec{a}_f - \omega \times (\omega \times r) - 2\omega \times v' - \omega \times v$$

$$F = m a' = m a_f + F_c + F_{Co} + F_g$$

$$F_c = -m \omega \times (\omega \times r)$$

$$F_{Co} = -2m \Omega \times v'$$

$$\omega = \frac{2\pi}{24 \cdot 3600} \hat{s}^{-1}$$

$$\Gamma_{ik} = -m \dot{x}_k$$


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17 Feb.

$$\ddot{x}_k = \begin{pmatrix} -1 & & & \\ & 0 & & \\ & & 1 & \\ & & & 1 \end{pmatrix} \ddot{x} = \ddot{x} \text{ if}$$

Flat spacetime

$$g_{ik} = \eta_{ik} a(t) + h_{ik}$$

$$= a(t) \eta_{ik} + h_{ik}(\vec{x}, t)$$

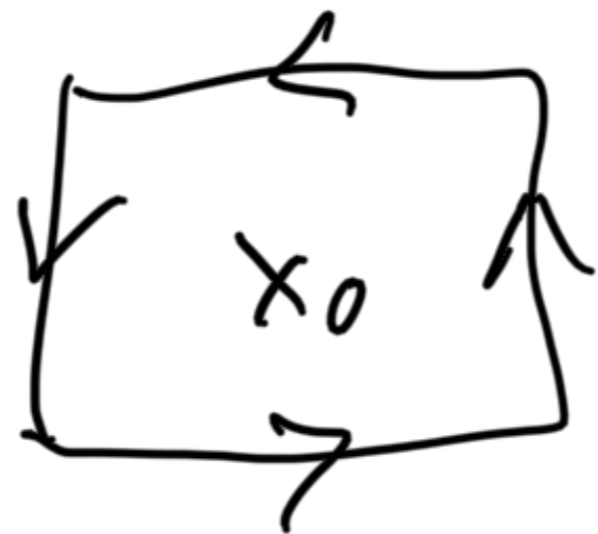
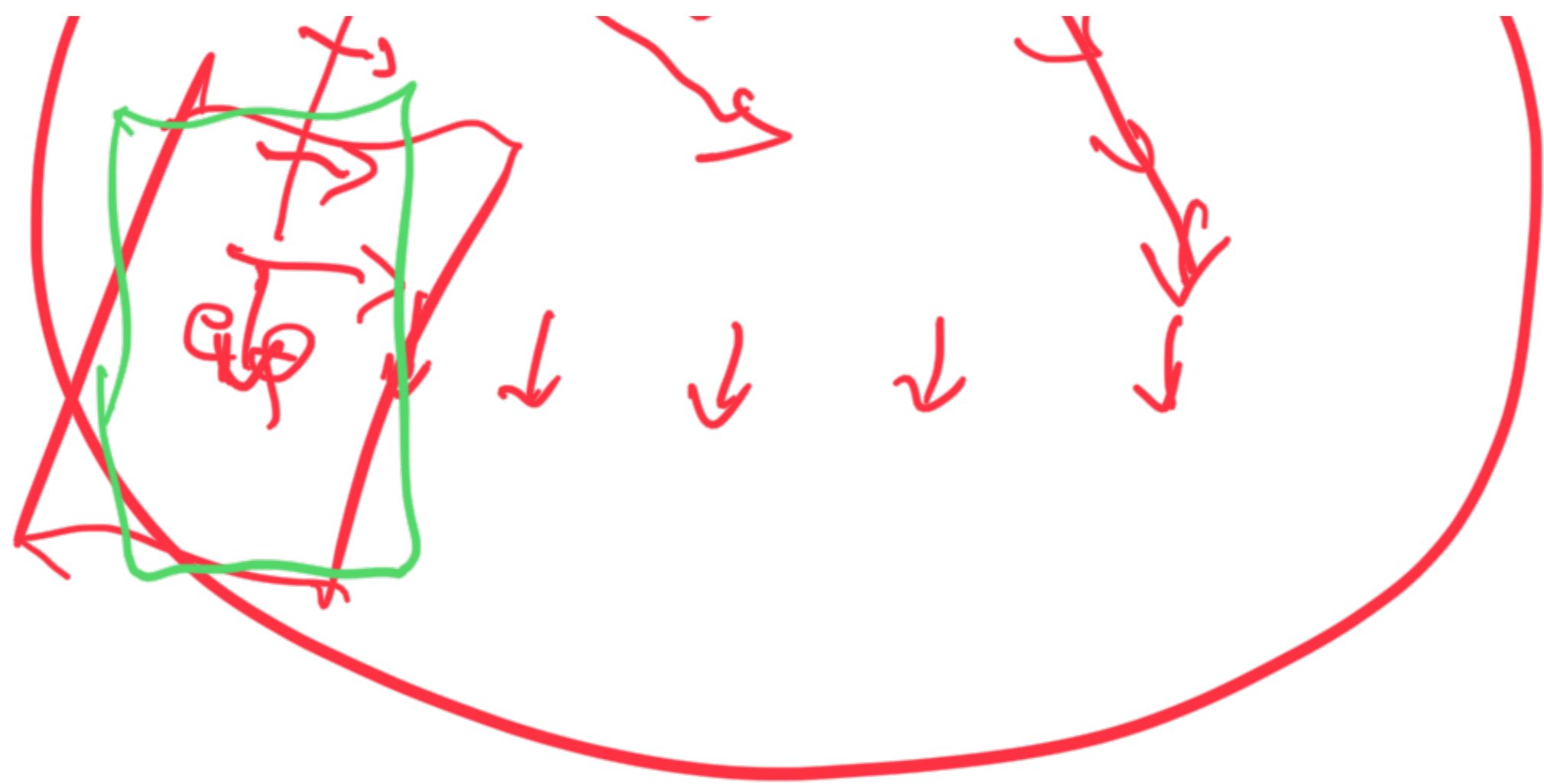
scale factor

$$a_0 = a(\text{now}) = a(3.8 \cdot 10^9 \text{ yrs}) = 1$$

$$a(t) \sim t^\alpha \quad \alpha = \frac{2}{3} \text{ Mat} \quad \frac{1}{2} \text{ rad}$$







$$\vec{1} = \begin{pmatrix} 1 \\ 1 \\ 1 \\ 1 \end{pmatrix}$$

$$R_{ie} = -\frac{1}{2} \square h_{ie} + \dots$$

$$\square = \nabla^2 - \partial_0^2$$

can remove by a

coordinate

transformation

$$\square = \partial^k \partial_k = \nabla^2 - \partial_0^2$$

$$x^{(i)} = x^{(i)} + \epsilon^{(i)}(x)$$

$$\begin{pmatrix} 2 \\ 1 \\ 0 \\ 1 \\ -2 \end{pmatrix}$$

$$\epsilon^{(i)} = 0$$

$$\begin{aligned}
 \vec{A} &= \vec{A} \parallel 0 \\
 \vec{B} &= \vec{B} \parallel 0 \\
 \vec{C} &= \vec{C} \parallel 0
 \end{aligned}$$

$$\vec{A} = \sum_{\vec{k}, \lambda} \left[ \epsilon_{\vec{k}}(x, \lambda) a(\vec{k}, \lambda) e^{-i\vec{k} \cdot \vec{x}} + \epsilon^*_{\vec{k}}(\vec{k}, \lambda) a^\dagger(\vec{k}, \lambda) e^{i\vec{k} \cdot \vec{x}} \right]$$

$$\vec{A} \parallel 0$$

$$H = \sum_{\vec{k}} \omega_k \left( a^\dagger(\vec{k}) a(\vec{k}) + \frac{1}{2} \right)$$

$$\langle 0|H|0\rangle = \frac{1}{2} \sum_{\vec{k}} \sqrt{\vec{k}^2 + m^2}$$

$$E = \langle 0|H|0\rangle$$

$$= \frac{1}{2} \int d^3k \sqrt{k^2 + m^2}$$

$$= 2\pi \int_0^{\infty} k^2 dk \sqrt{k^2 + m^2}$$

$$\Rightarrow 2\pi \frac{\infty^4}{4} = \frac{\pi}{2} (\infty)^4 \text{ (*)}$$

$$\Rightarrow \frac{\pi}{2} M_p^4 \sim (10^{19})^4$$

$$\langle H \rangle = (10^{13} \text{ eV})^4$$

4

22 4

$$\left( \frac{10^{19}}{10^{-3}} \right)$$

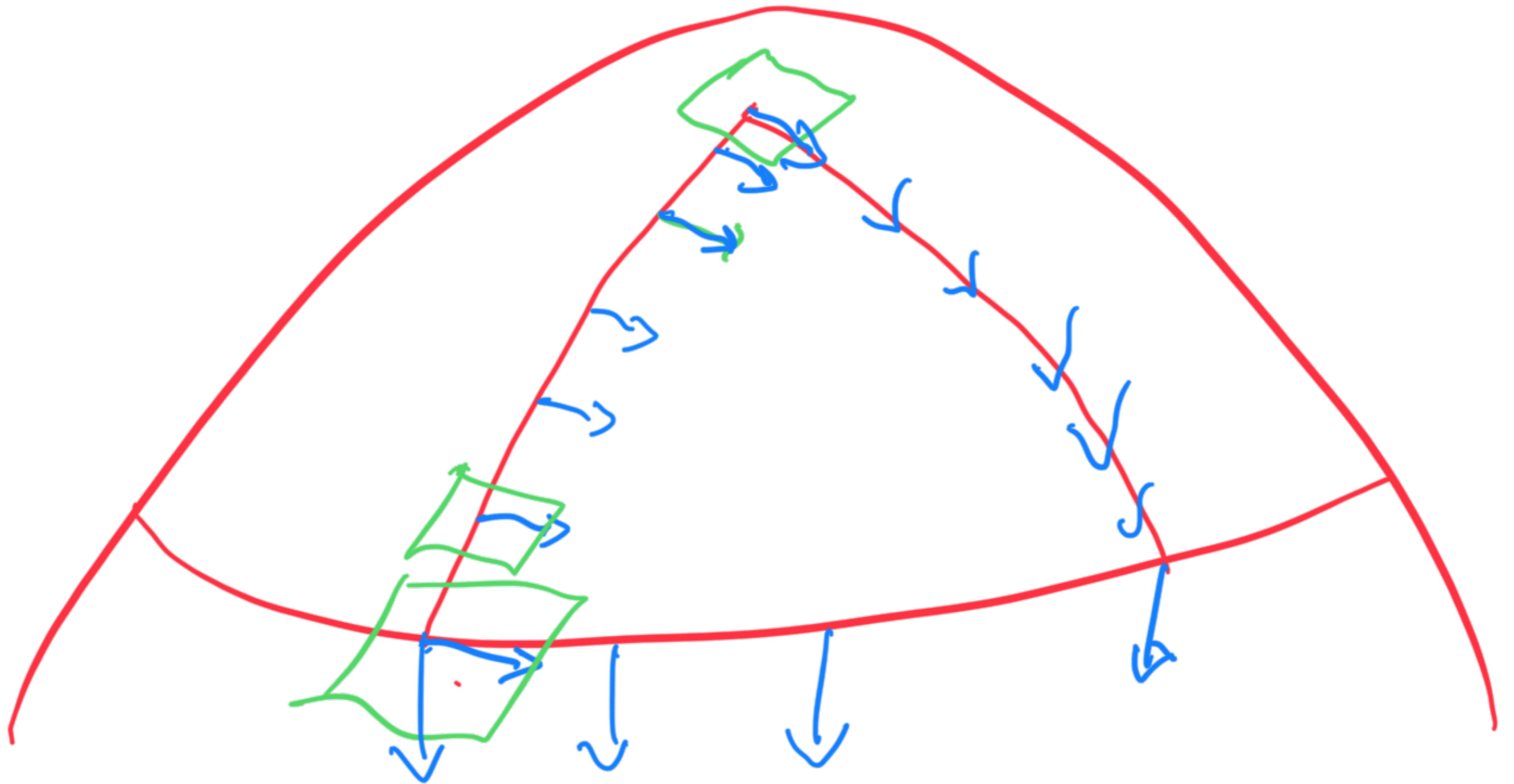
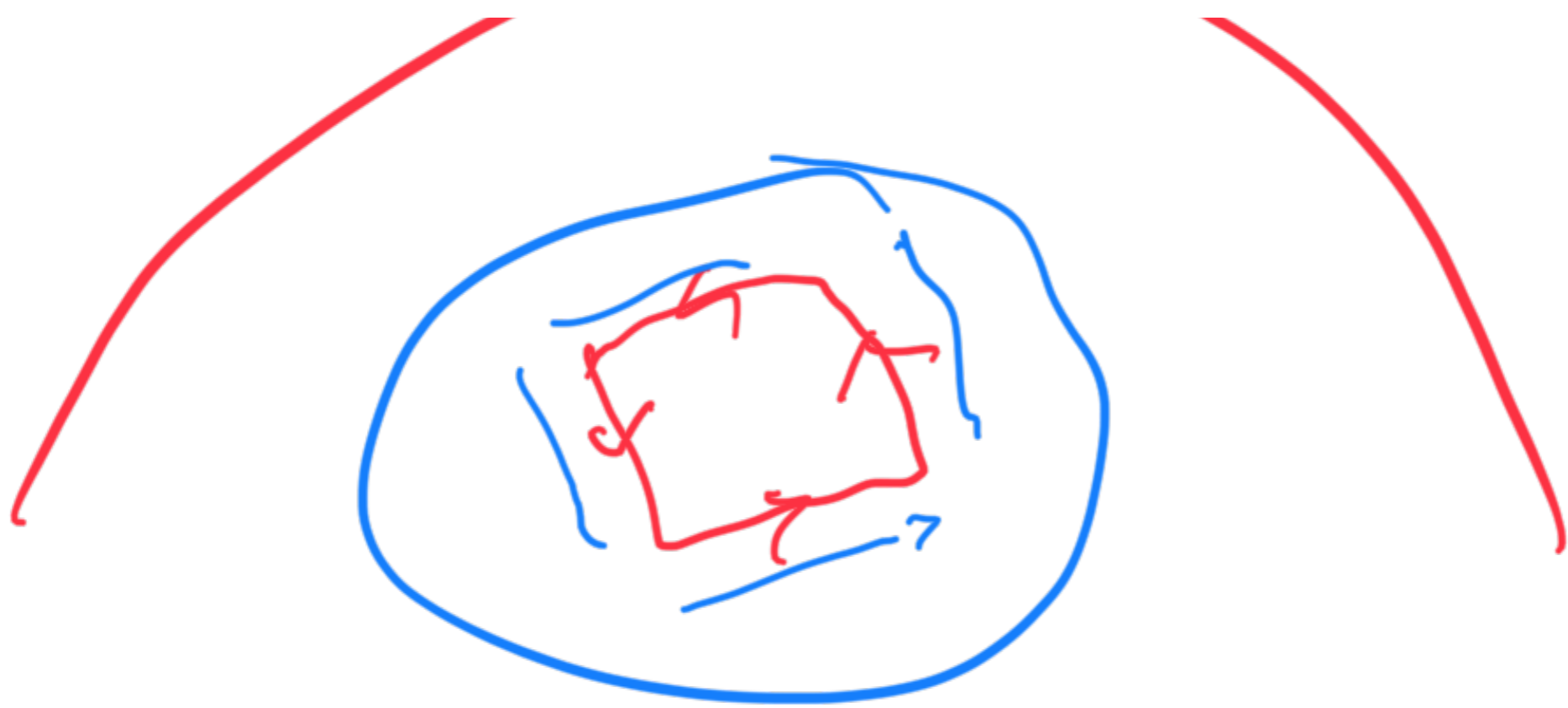
$$\begin{aligned} & 17 \quad 10^{22.7} \\ & = 10^{88} \end{aligned}$$

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FEB 2022

$$S = \frac{c^3}{16\pi G} \int (R + \Lambda) \sqrt{g} \, d^4x$$

cosmological constant





$$R = R^i_{\cdot}$$

$$S \propto \int_R \sqrt{g} d^4x$$

$$\propto \int (R + \Lambda) \sqrt{g} d^4x$$

$$\propto \int |\det(g_{ik})|$$

$\wedge$   $\equiv$  ?  $\equiv$  dark energy

1

Peebles

$$w \sum_i \frac{p_i}{c^2 \rho_i}$$

Sternhardt  
Carroll

quintessence

$$1 \Leftrightarrow w = -1$$

$$[a(p), a^{\dagger}(k)] = \delta^3(\vec{p} - \vec{k})$$

$$= \int \frac{d^3x}{(2\pi)^3} \delta^3(\vec{p} - \vec{k})$$

$$\int d^3x = V$$

$$[a(p), a'(p)] = 0 \quad (2\pi)^3 \quad (2\pi)^3$$

$$\int d^3p \sqrt{p^2 + m^2}$$

$$\approx 4\pi \int_0^{\infty} dp \, p^2 \sqrt{p^2 + m^2}$$

quantitatively

$$\int_0^{\infty} dp \, p^3 = 4\pi \frac{p^4}{4}$$

divergent energy

-30

$$\rho \approx 7 \times 10^9 \text{ cm}^{-3} = 7 \times 10^{27} \text{ kg m}^{-3}$$

$$\sim (2 \times 10^5 \text{ eV})$$

$$\hbar = c = 1$$

$$10^3 \text{ eV} \sim m_\nu$$

$$m_{\nu 1}^2 - m_{\nu 2}^2 =$$

$$m_{\nu 2}^2 - m_{\nu 3}^2 =$$

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With then

exact SUSY  $\Rightarrow E_0 = 0$ .

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string theory

$\checkmark$

$\phi$

AVOIDS

$$[\phi(x,t), \phi(y,t)] = i \delta^3(\vec{x} - \vec{y})$$

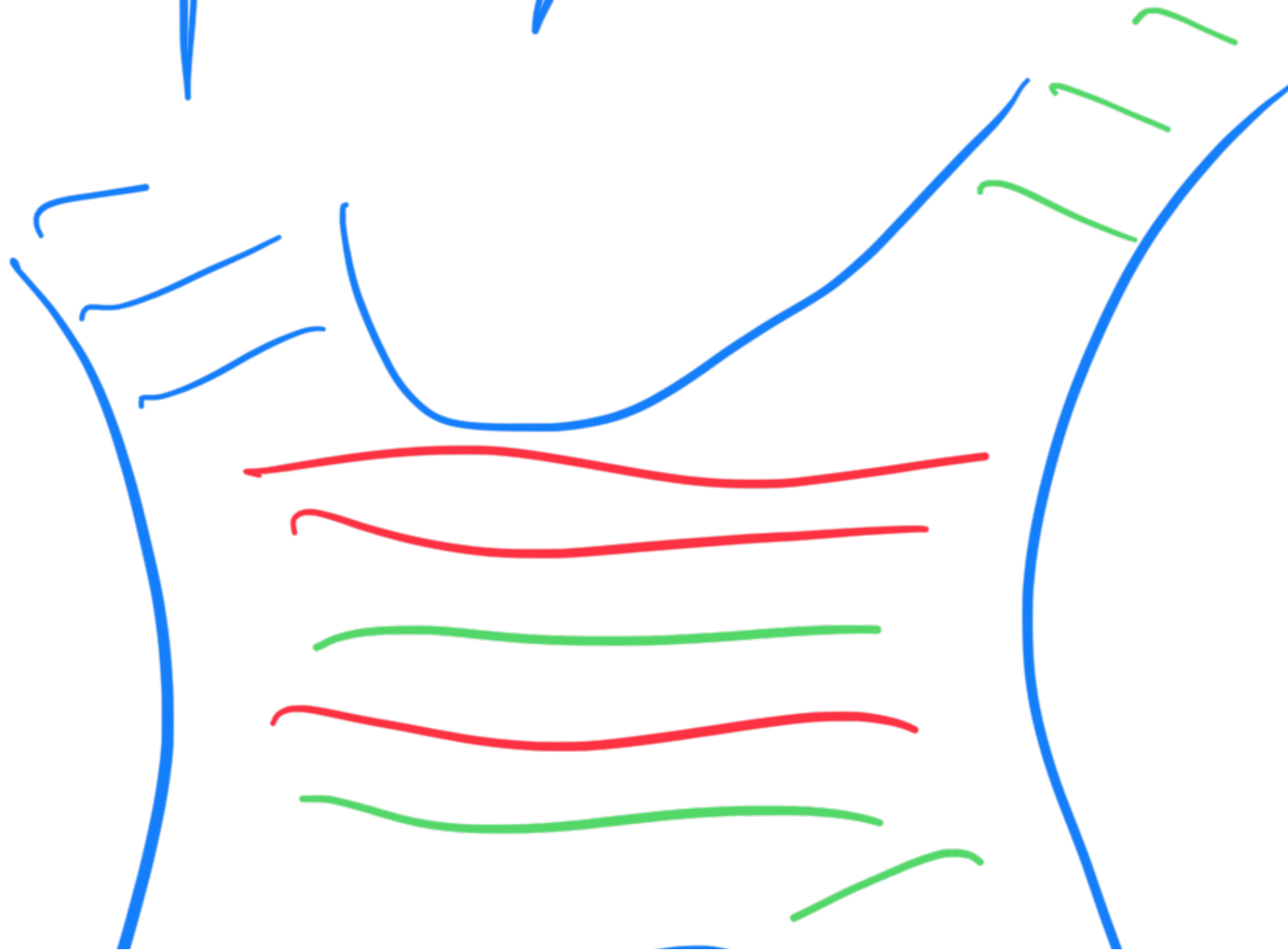


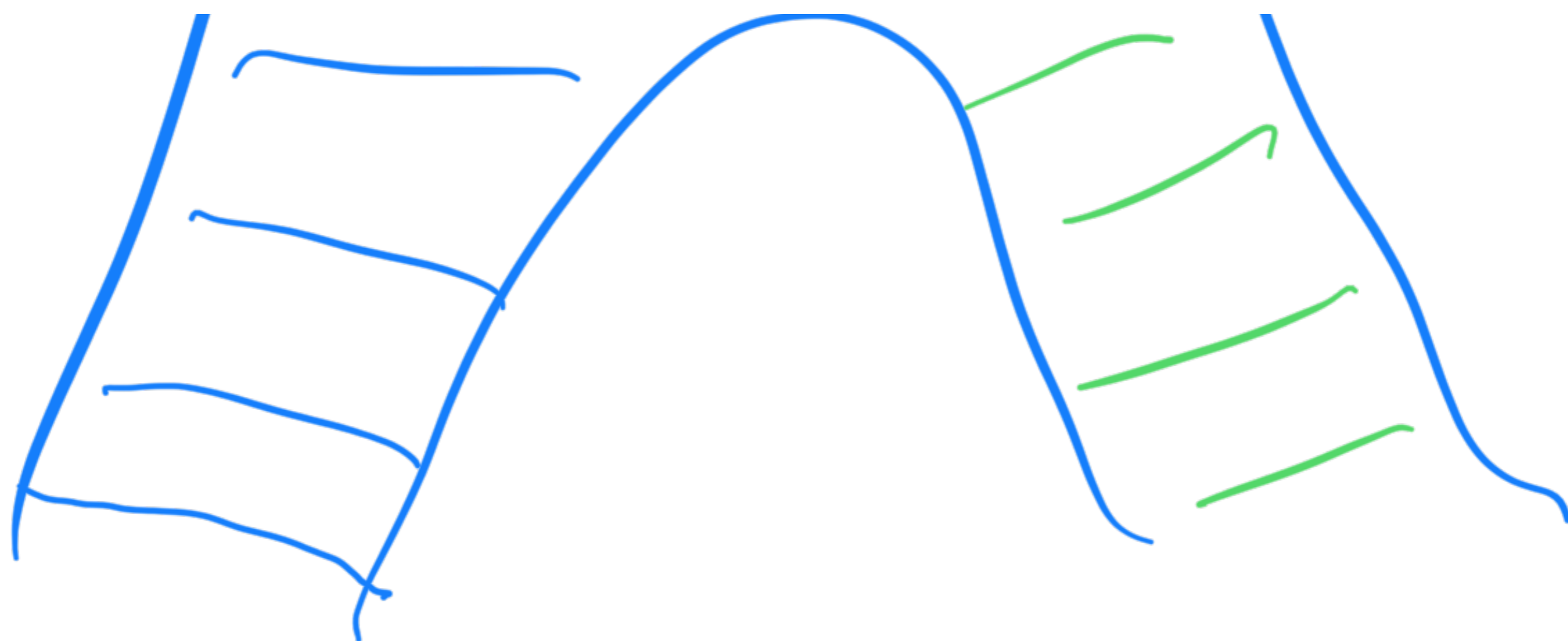
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21 March

$$a_k^\dagger a_p + a_p^\dagger a_k = \delta(p-k)$$

$$E^0 = \sqrt{m^2 + \vec{p}^2}$$





$$H \approx \omega \left( a^\dagger a + \frac{1}{2} \right)$$

$$a = \frac{q + ip}{\sqrt{2}}$$

$$\left( \dots \left( \frac{q + ip}{\sqrt{2}} \right) \dots \right)$$

$$H = \omega \left( \frac{p^2}{2m} + \frac{1}{2} m \omega^2 x^2 \right)$$

$$= \frac{\omega}{2} (p^2 + m^2 \omega^2 x^2) + i \frac{\omega}{2} [x, p^2]$$

$$= \frac{\omega}{2} (p^2 + m^2 \omega^2 x^2) - \frac{\omega}{2}$$

|| >

-



$= \text{wata}$

$$\langle g | H | 0 \rangle = 0$$

$$\langle g | (g + ip) | 0 \rangle = 0$$

$$\rightarrow \langle g | \phi \rangle + i \langle g | p | 0 \rangle = 0$$

$$\rightarrow \langle g | 0 \rangle + \frac{\partial}{\partial g} \langle g | 0 \rangle = 0$$

$$\frac{d}{dg} \langle g | 0 \rangle = -g \langle g | 0 \rangle - g^2/2$$

$$\langle g | 0 \rangle = 0$$


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$$H = (W - ip)(W + ip)$$


---

$$\frac{d}{dg} \langle g | 0 \rangle = -W \langle g | 0 \rangle$$

$$\langle 10 | 0 \rangle = 0$$

$$= 0$$

Weg' d'g'

$$| \langle 10 | 0 \rangle | \rightarrow 0$$

$$\rightarrow 0$$

as  $\uparrow \rightarrow 0$



6'3



5'

11



4

4



4

1 1 1 1 1 1 1 1

1 2 3 4 5 6

0 1 2 3 4 5 6

1 2 3 4 5 6

1 2 3 4 5 6



1 2 3 4 5 6

$$\begin{aligned}
 H &\approx \int d^4x (W - i\pi)(W + i\pi)(x) \\
 &= \int d^4x (W - i\pi) \left( \sqrt{p^2 + m^2} + i\phi + i\pi \right)
 \end{aligned}$$

$$\pi(x) \rightarrow \frac{1}{i} \frac{\delta}{\delta \phi(x)}$$

$$\langle \phi | (W + i\pi) | \Omega \rangle = 0$$

$$\dots \langle \phi | \Omega \rangle = 0$$

$$\approx W(\phi) \langle \phi | \Omega \rangle + \frac{\delta W(\phi)}{\delta \phi}$$

$$\frac{\delta}{\delta \phi} \langle \phi | \Omega \rangle = - W \langle \phi | \Omega \rangle$$

$$\int dx \, n (f + \epsilon h)^{n-1} h$$

$$\rightarrow \int n f^{n-1} h dx$$

$$G = \int dx f^n(x)$$

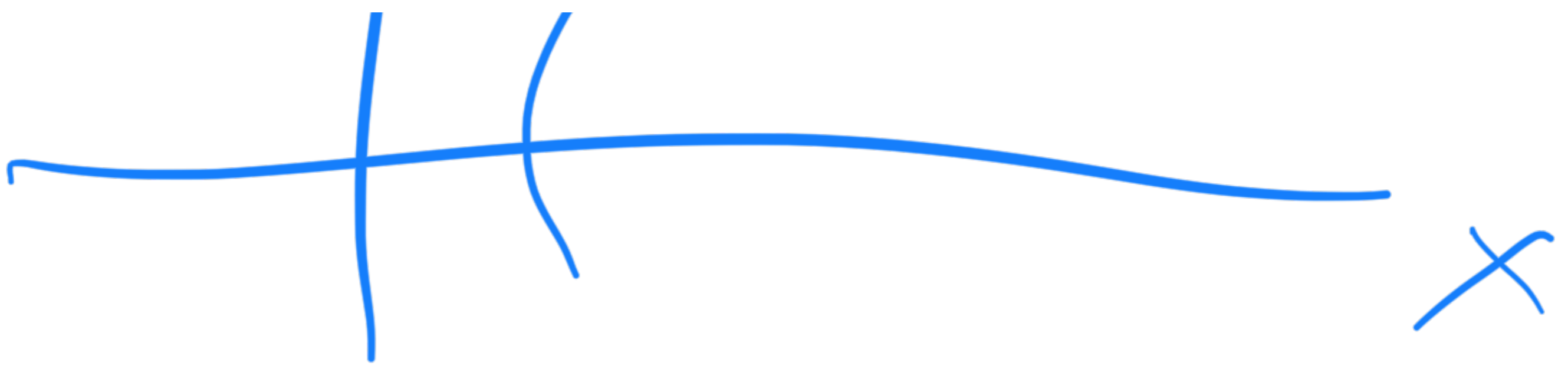
$$\delta G = \delta G [f, \delta_y]$$

$$\delta f(y)$$

$$\delta_y(x) = \delta(x-y)$$

y





$$\sqrt{dx^2 + dy^2} = dx \sqrt{1 + \frac{dy^2}{dx^2}}$$

$$= dx \sqrt{1 + \left(\frac{dy}{dx}\right)^2}$$